

Figure 1: Raw count rate curves of GRB 230126A recorded by the detector unit no. 0. The vertical dashed line marks the trigger time by Fermi/GBM.

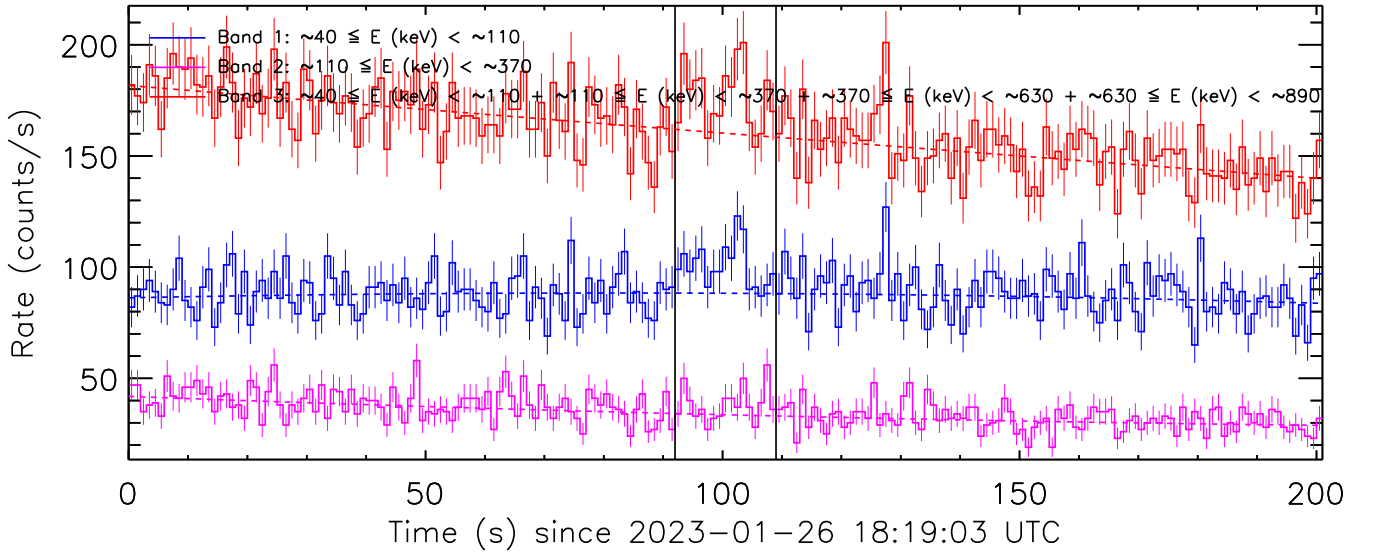


Figure 2: Background fit to the data from the detector unit no. 0 by a second order polynomial. Pre-burst and post-burst intervals which were used for the background fit are delimited by the vertical solid lines.

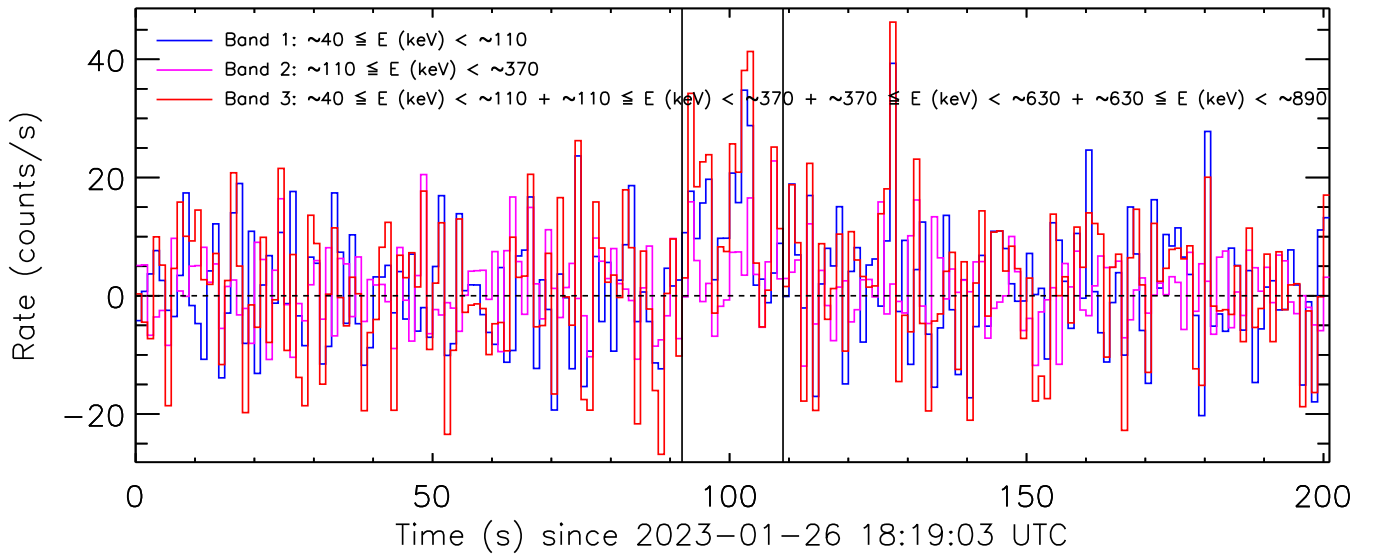


Figure 3: Background-subtracted count rate curves from the detector unit no. 0.